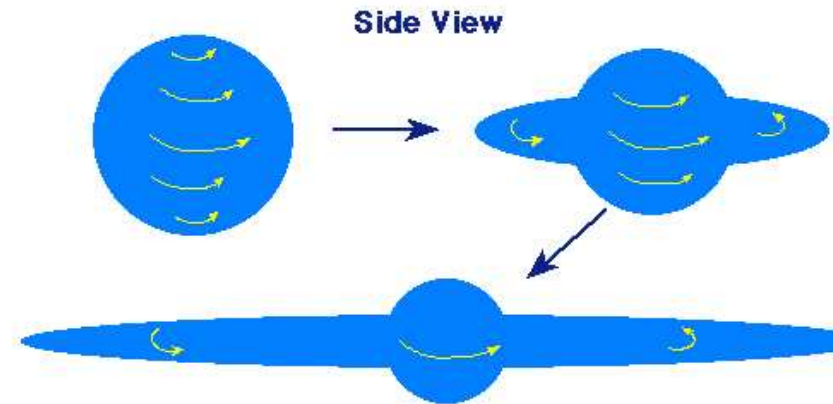


# Star Formation

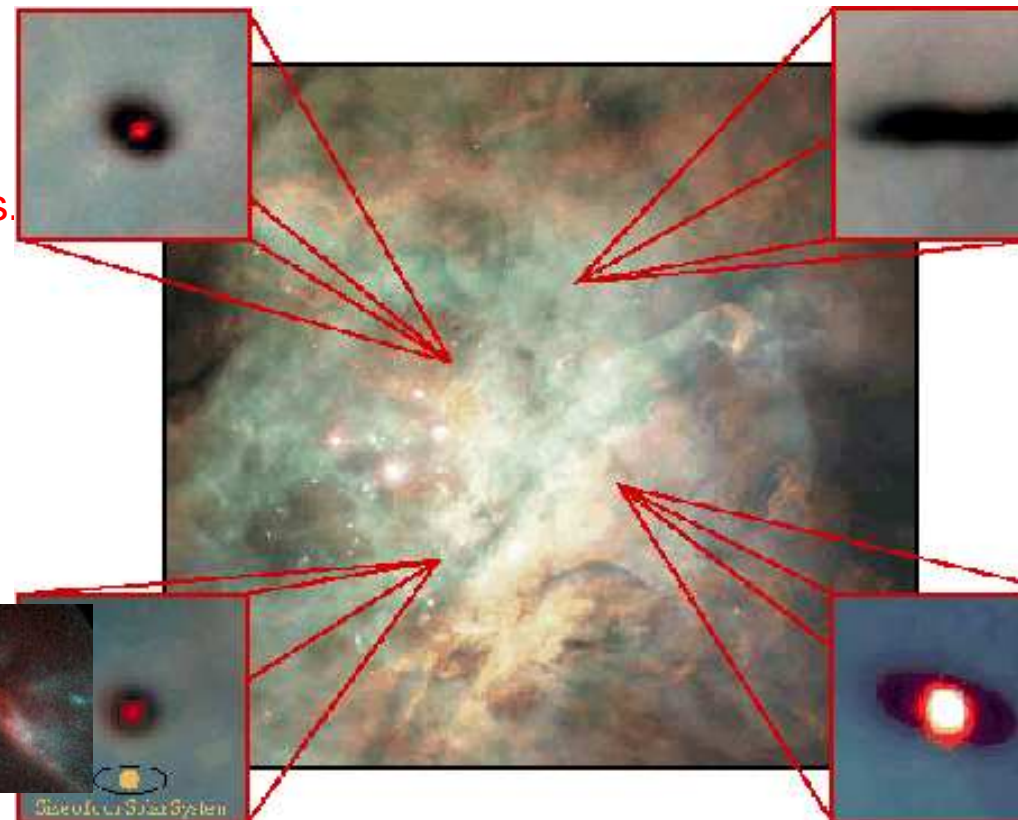
- Dense cores of molecular clouds collapse into hot plasma which eventually triggers nuclear reactions.
- Release of gravitational energy both heats the material and produces infrared radiation.
- Conservation of angular momentum requires spin rate to increase during collapse and disc formation.
- Emissions of protostar hidden by cooler infalling matter; eventually breaks out at poles (least resistance), producing jets.
- Dense clumps form in protostellar disc that eventually themselves collapse into planets.
- Sufficiently rapidly rotating collapses form double or multiple stars.
- Young star's emissions sweep out gas and dust from planetary system.



[csep10.phys.utk.edu/astr161/lect/solarsys/nebular.html](http://csep10.phys.utk.edu/astr161/lect/solarsys/nebular.html)



B. Reipurth (Univ. of Col.), NASA



C. R. O'Dell and S. K. Wong (Rice U.), WFPC2, HST, NASA

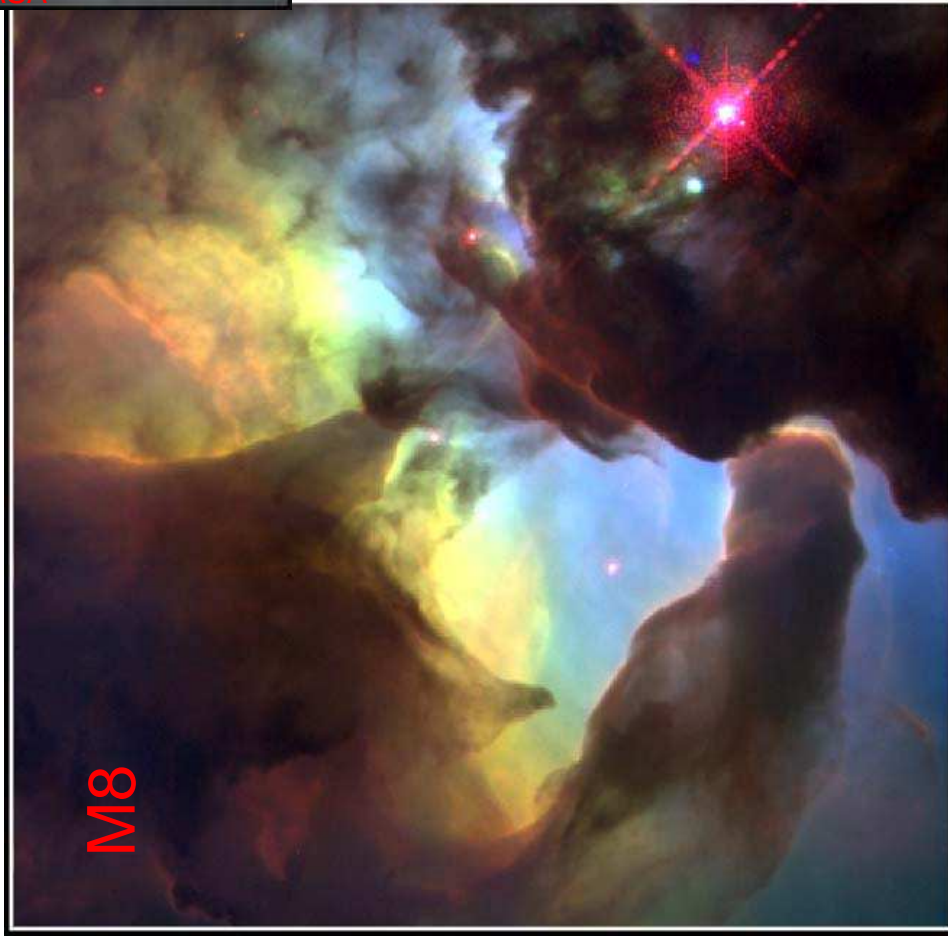
# Star Forming Regions

M42



C.R. O'Dell (Rice U.), NASA

M8



Lagoon Nebula Detail

PRC96-38b • ST ScI OPO • January 22, 1997  
A. Caulet (European Southern Observatory) and NASA

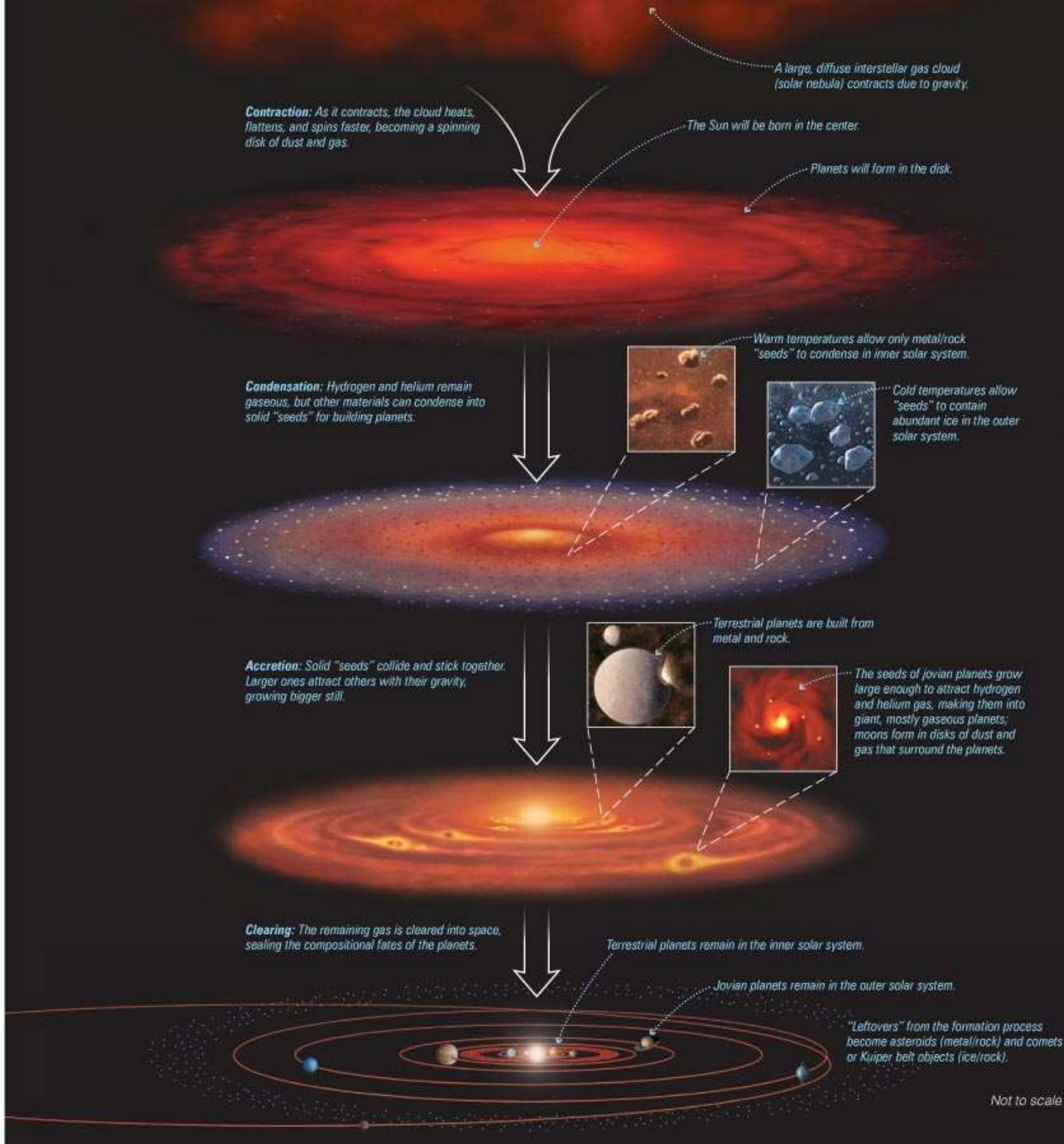
HST • WFPC2

H. Yang (UIUC) & J. Hester (ASU), NASA

NGC 604

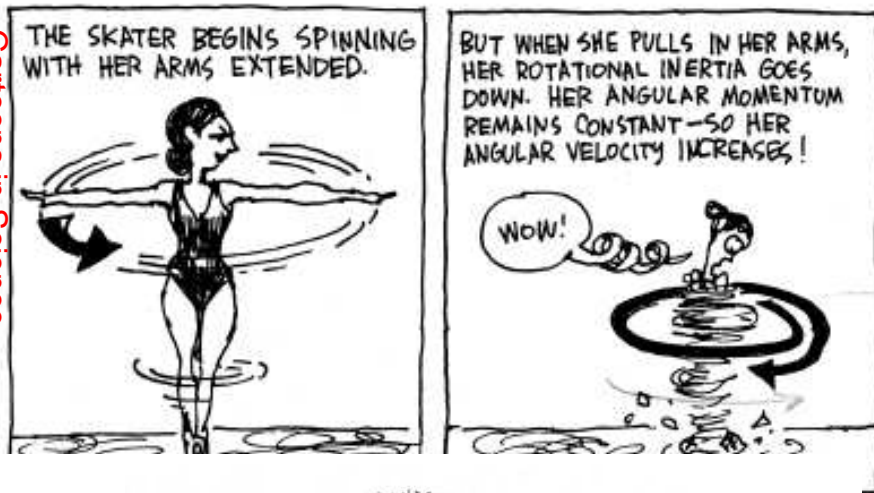










# Angular Momentum and Condensation

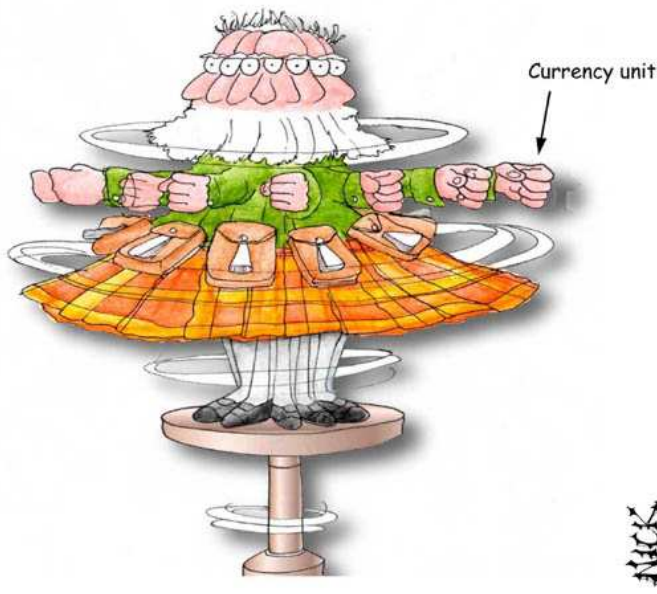
REMEMBER THAT MOMENTUM IS CONSERVED IN THE ABSENCE OF EXTERNAL FORCES. LIKEWISE, **ANGULAR** MOMENTUM IS CONSERVED IN THE ABSENCE OF EXTERNAL **TORQUES**.



Materials in the Solar Nebula				
	Metals	Rocks	Hydrogen Compounds	Light Gases
Examples	 iron, nickel, aluminum	 silicates	 water (H <sub>2</sub> O) methane (CH <sub>4</sub> ) ammonia (NH <sub>3</sub> )	 hydrogen, helium
Typical Condensation Temperature	1,000–1,600 K	500–1,300 K	< 500 K	(do not condense in nebula)
Relative Abundance (by mass)	■ (0.2%)	■ (0.4%)	■ (1.4%)	■ (98%)

Addison-Wesley

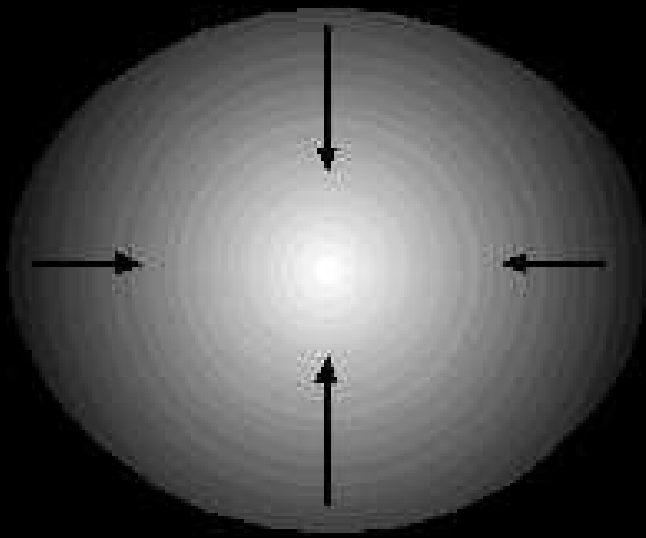
Nick D. Kim, nearingzero.net



NICK

## Centrifugal Force, $n$ .

The strong inward force experienced by a unit of currency clasped in the hand of a rotating Scotsman.

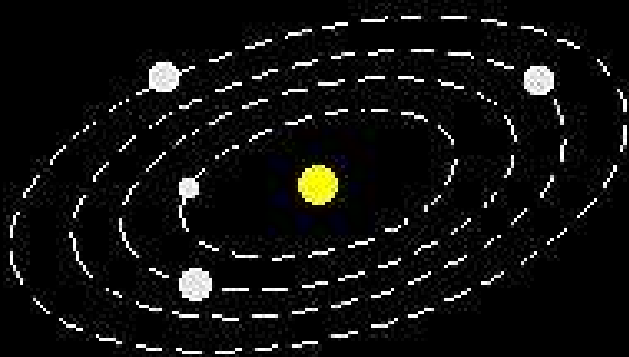


1) Large gas cloud collapses. Its spin causes it to flatten into a disk.

2) *side-view of inner part of disk:* warmer, denser, and more material closer to hot proto-Sun.



2) *above flat disk:* at an angle rock and metals condense all over the nebula; water, methane, ammonia, other ices condense only in outer part of nebula. Most of the material accumulates at the center.



4) Sun begins fusion. Strong winds clear away nebula. Remaining planetesimals near planets get swept up or flung out.



3) Solid particles collide and stick together to make planetesimals. Jovians pull in hydrogen and helium gas (proto-Sun continues to do so).

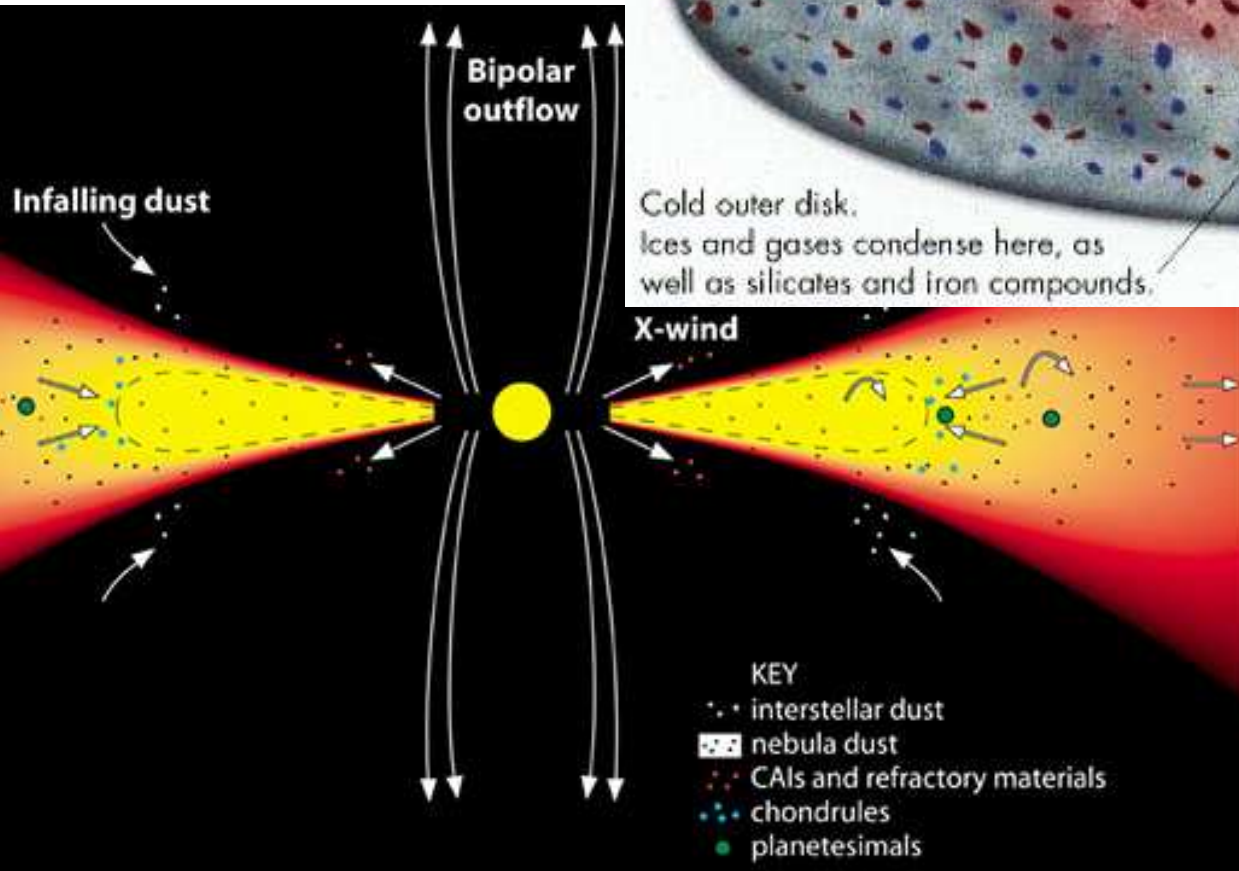
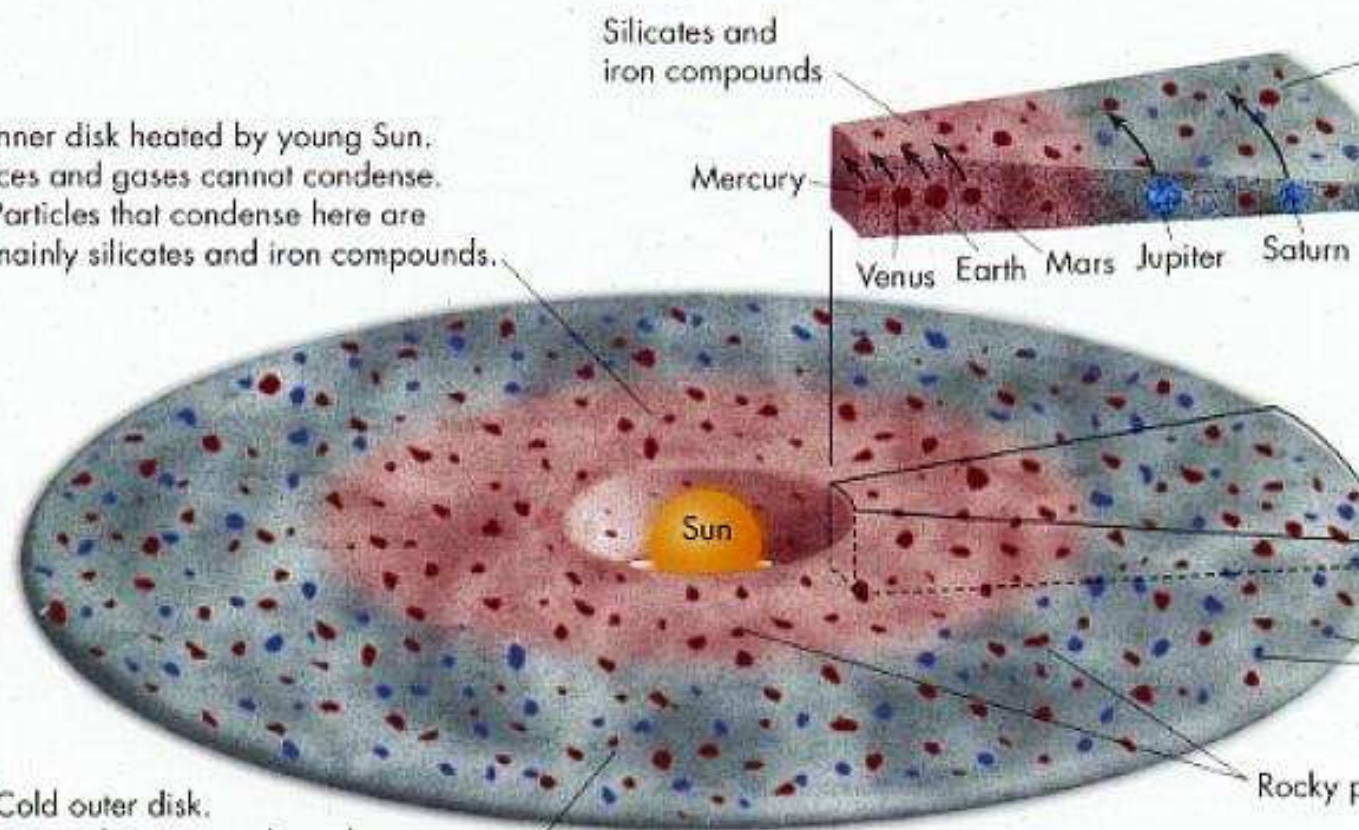
Condensation Model for making the solar system starting from upper left and proceeding clock wise (not all planets are shown).



# Planetismals

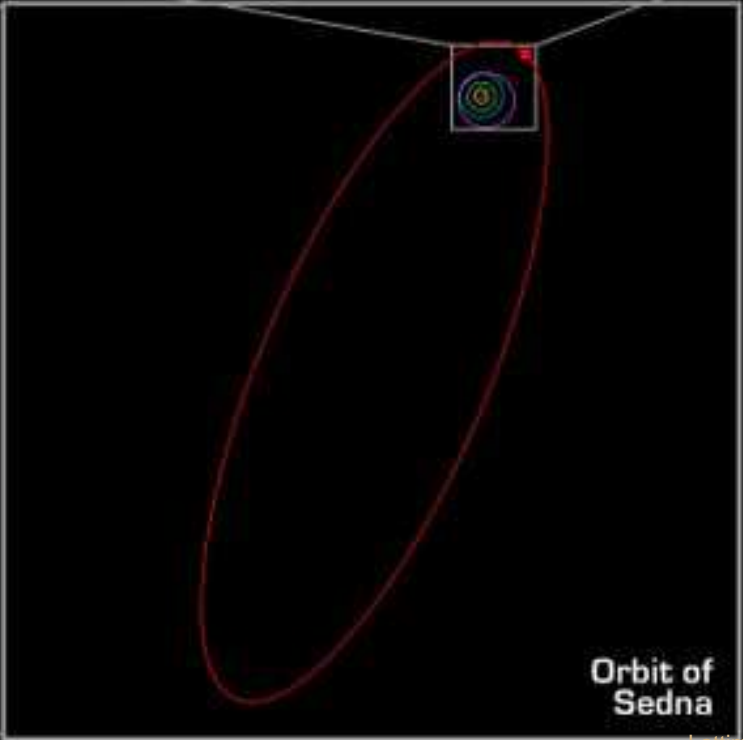
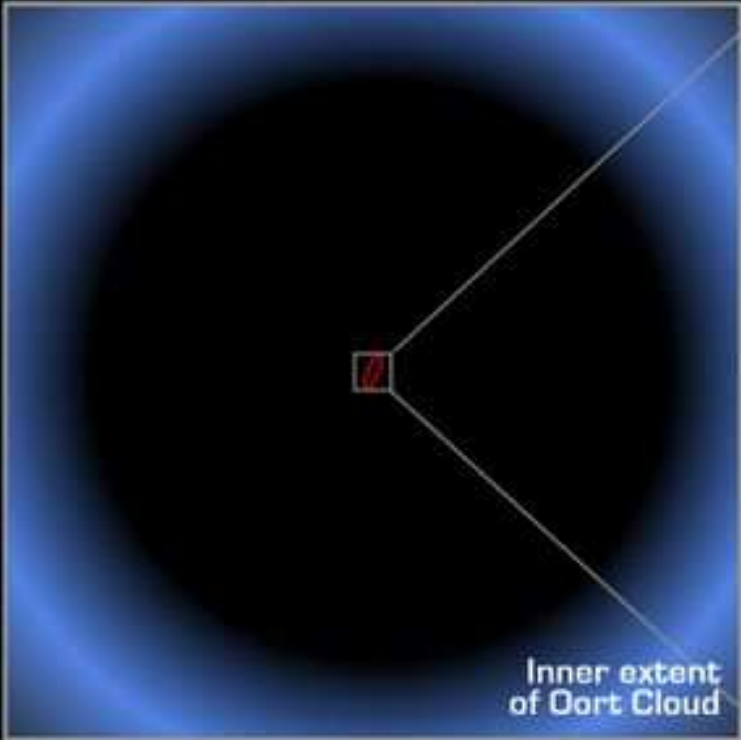
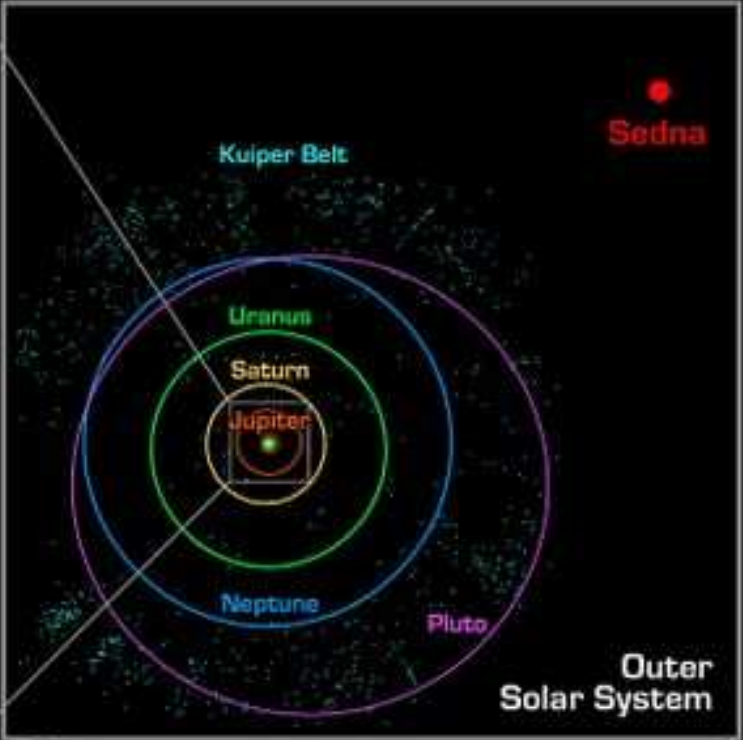
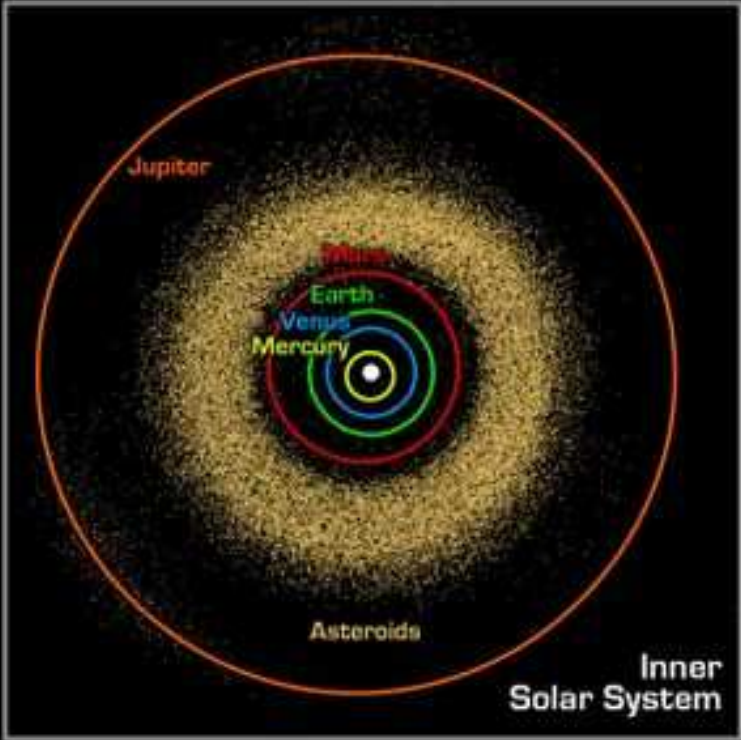
## Form

Inner disk heated by young Sun. Ices and gases cannot condense. Particles that condense here are mainly silicates and iron compounds.

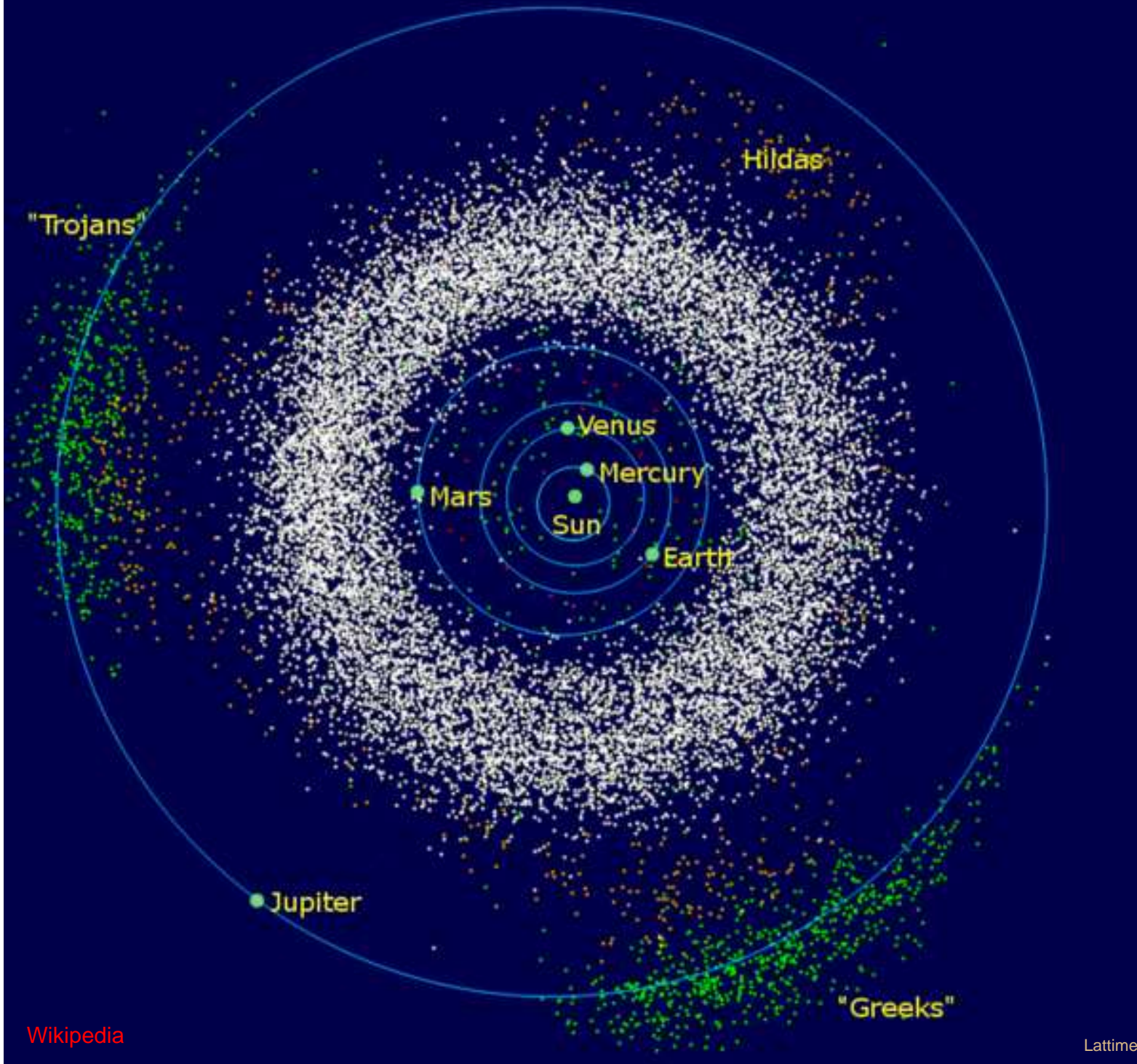


[www.astro.virginia.edu/class/skrutskie/astr121/notes/sscond.html](http://www.astro.virginia.edu/class/skrutskie/astr121/notes/sscond.html)

(PSRD graphic by Nancy Hurbirt, based on a conceptual drawing by Edward Scott, Univ. of Hawaii.)

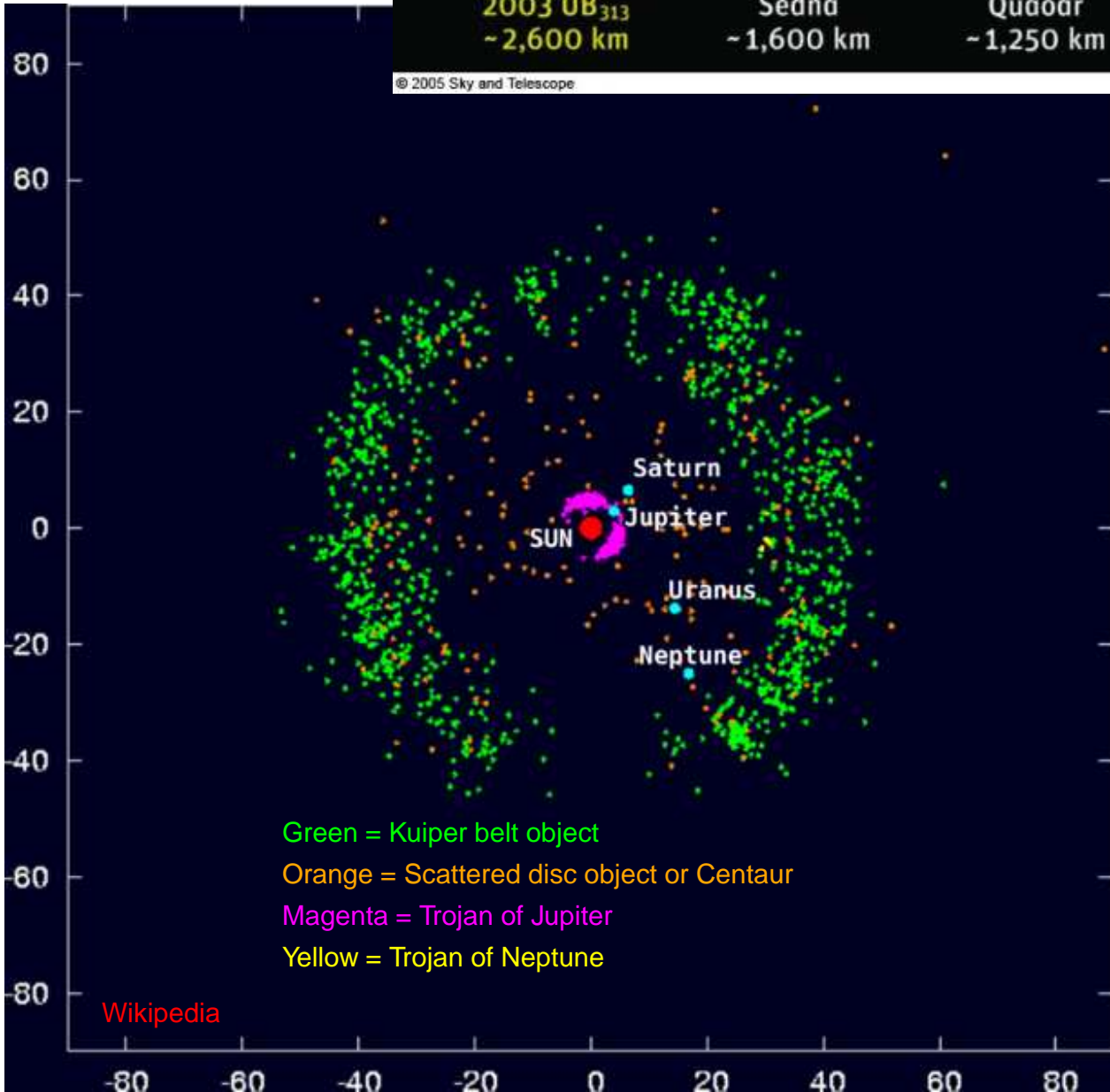








# Kuiper Belt

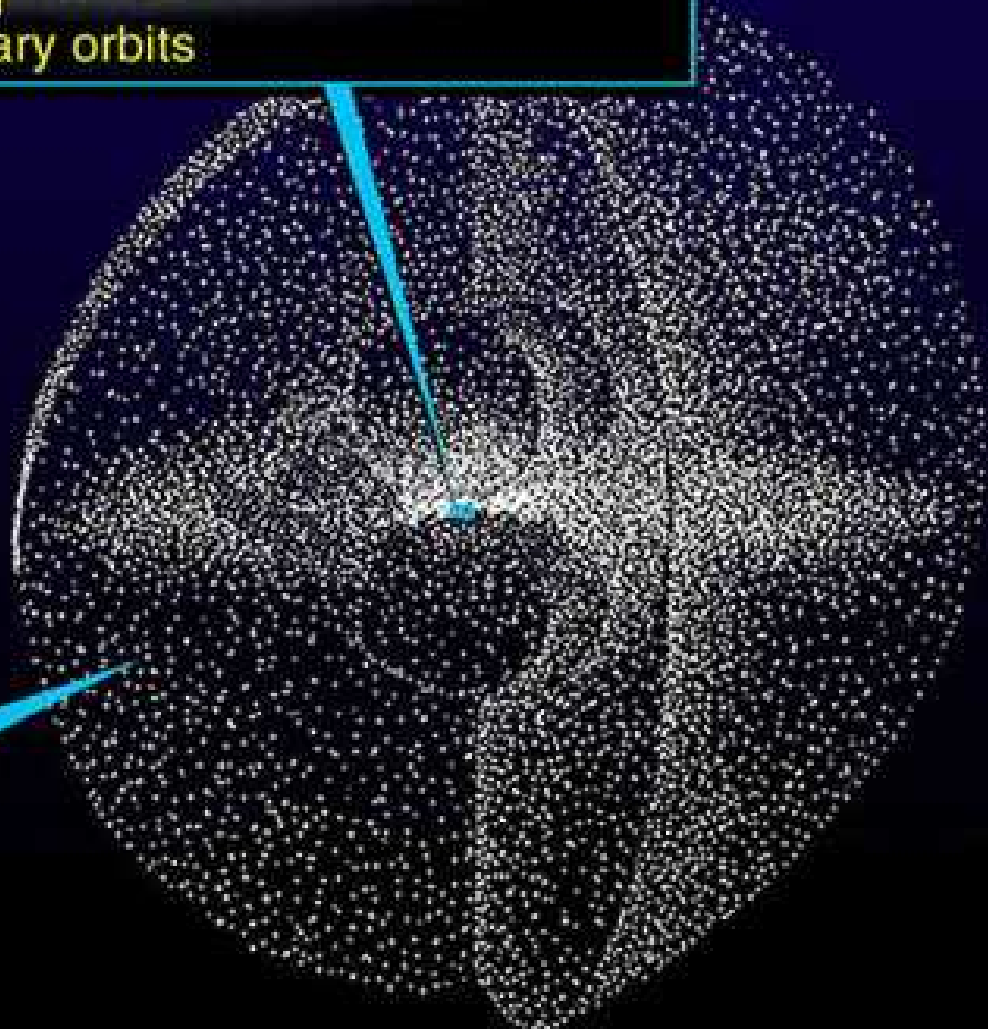


# Oort Cloud

[herschel.jpl.nasa.gov/solarSystem.shtml](http://herschel.jpl.nasa.gov/solarSystem.shtml)



The Oort Cloud  
(comprising many billions of comets)



*Oort Cloud cutaway drawing adapted from Donald K. Yeoman's illustration (NASA, JPL)*

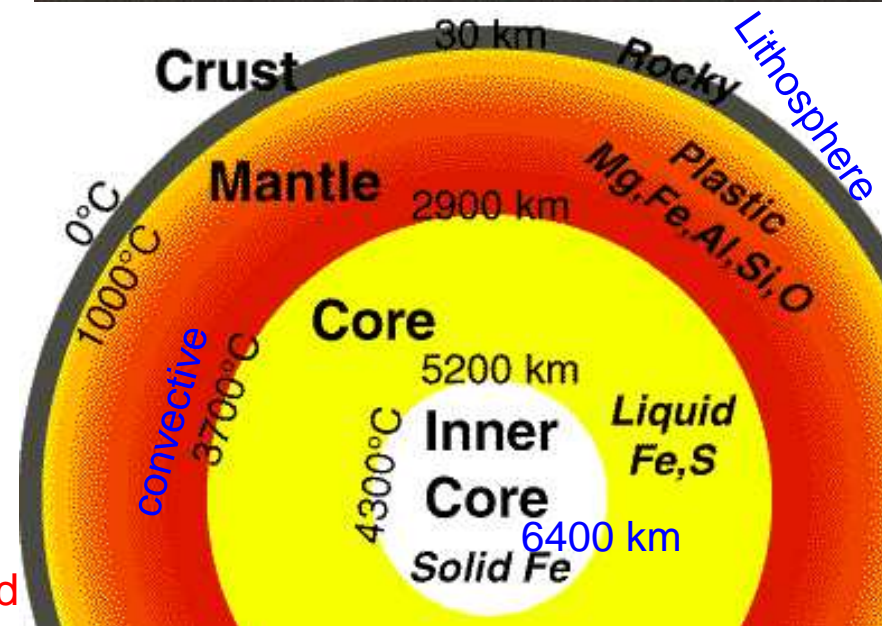


# Solar System



# Formation of the Earth

- Earth formed 4.5 Gyrs ago.
- From a small accumulation of gas and dust, it grew by accretion and bombardment of rocky material.
- Impacts and radioactive energy release kept the early Earth very hot and molten.
- A large impact about 4.4 Gyr ago resulted in the formation of the Moon.
- The moon is the largest natural satellite of the Earth; the other moons being Cruithne , asteroid 2003 YN107, 1998 UP1 and 2000 PH5.
- The peak “Late Heavy Bombardment” was from about 4.0-3.8 Gyrs ago.
- The high temperature of the early Earth promoted differentiation, which resulted in the formation of the Earth’s core and mantle as heavier elements, like Fe, Ni, Co, Mn and S sank and lighter elements such as Si, Mg and O floated.
- The molten character of the Earth destroyed most traces, but some 4.4 Gyr-old zircons have survived
- The Earth’s crust solidified and oceans formed about 4 Gyr ago.
- Life seems to have formed on the Earth about the same time.



[www.physci.wsc.ma.edu/young/hgeol/geoim/timeline/formation/formation.html](http://www.physci.wsc.ma.edu/young/hgeol/geoim/timeline/formation/formation.html)